

A CO Survey of Galactic Stellar Clusters

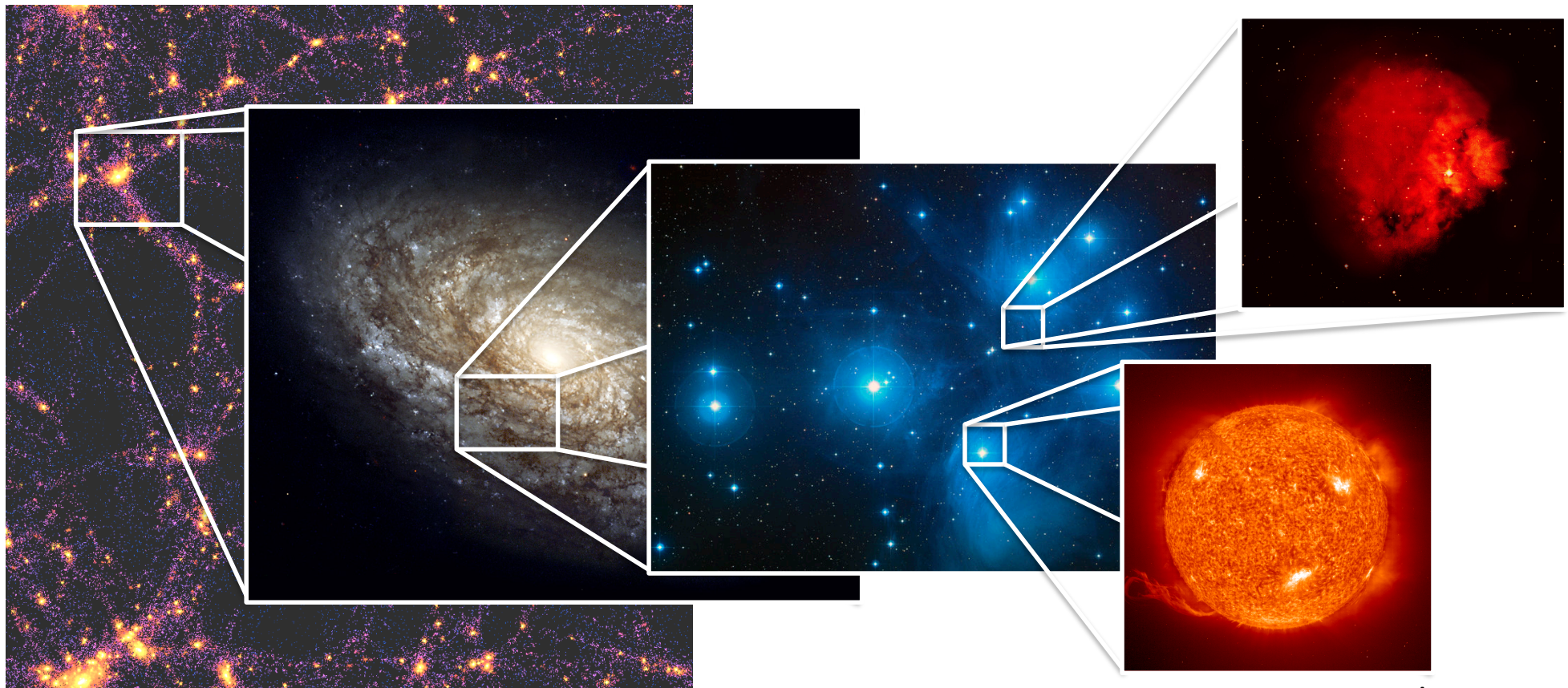
In the

Second Galactic Quadrant

Credit ; Eli Bressert

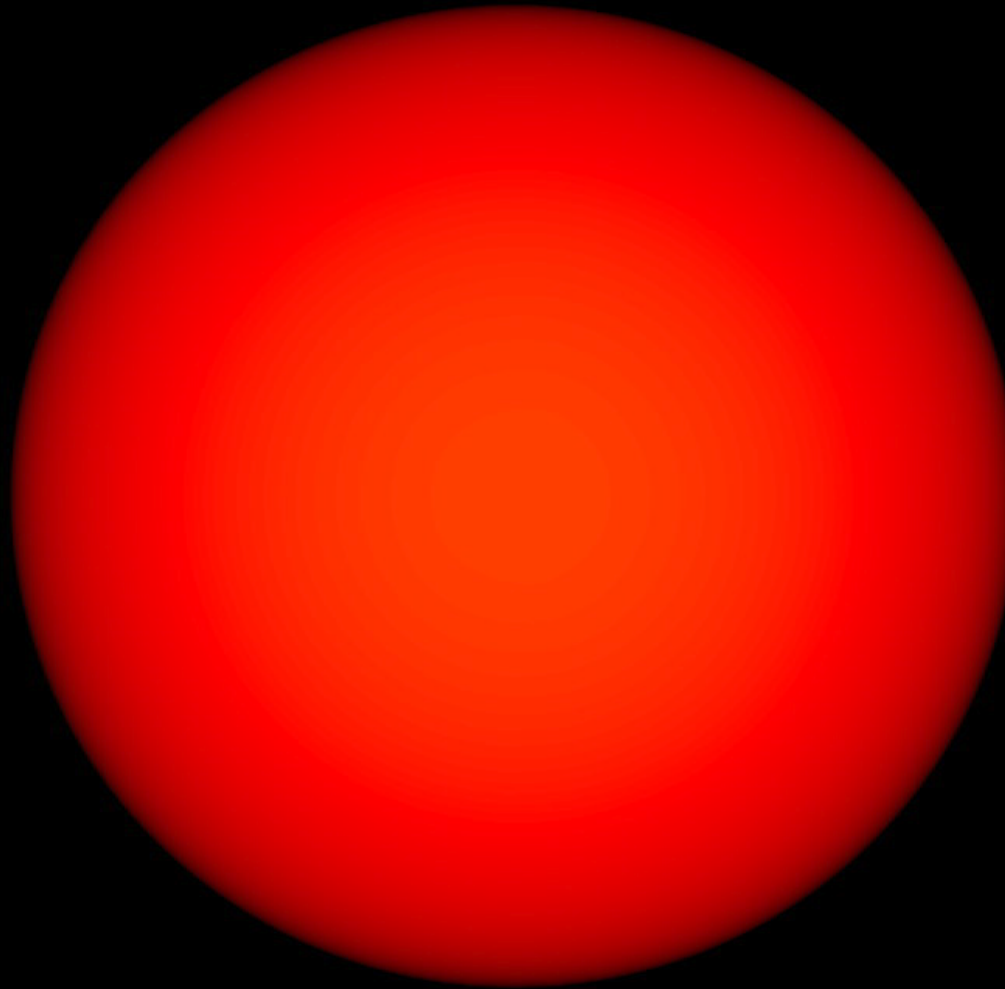
October 2009

Star Formation

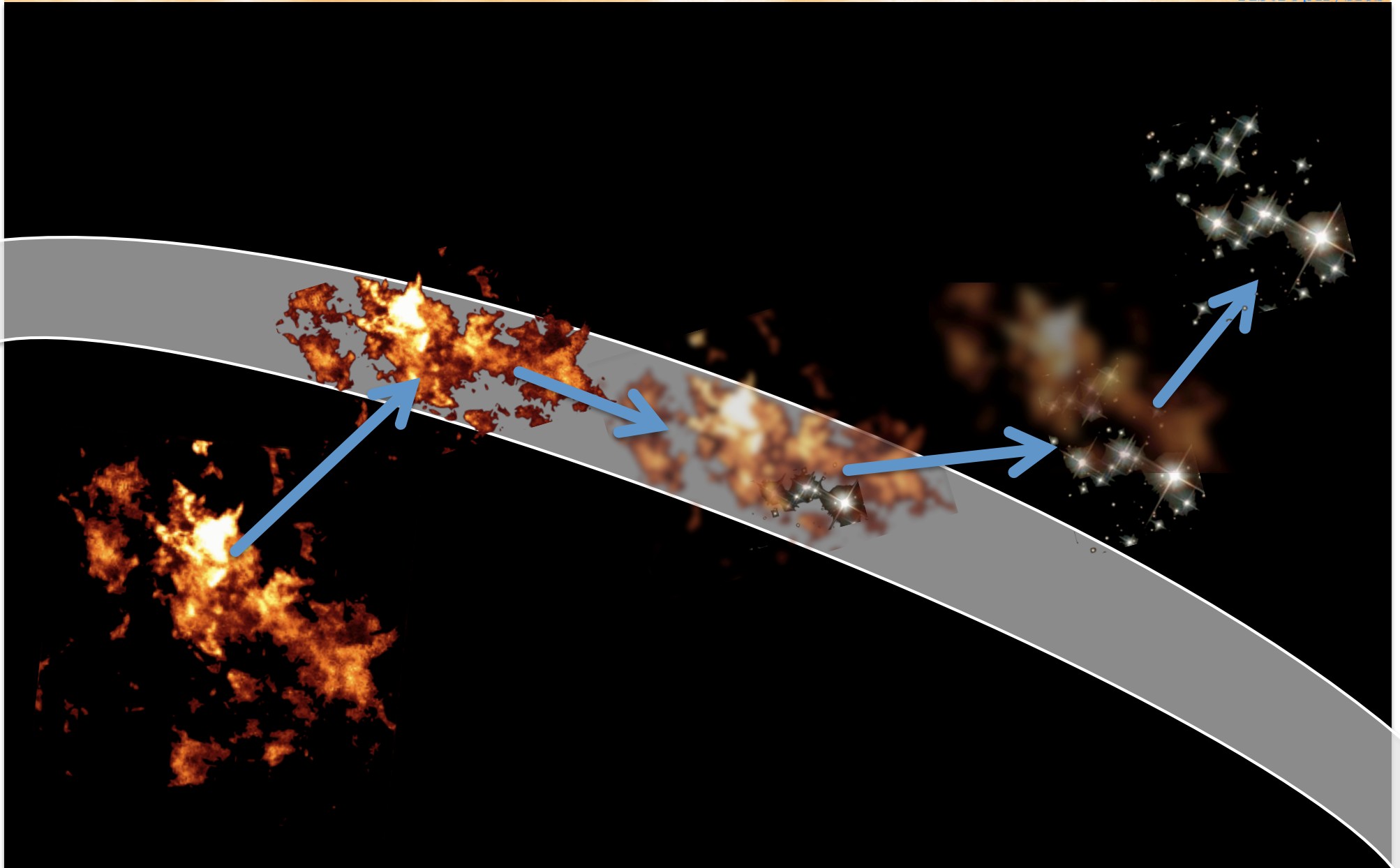


Star Formation





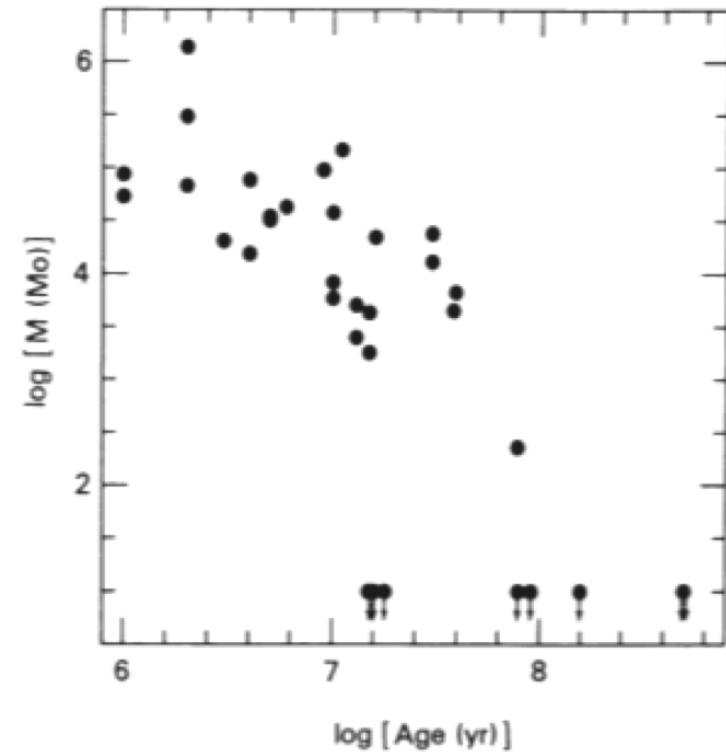
Star Formation



Lee Summers

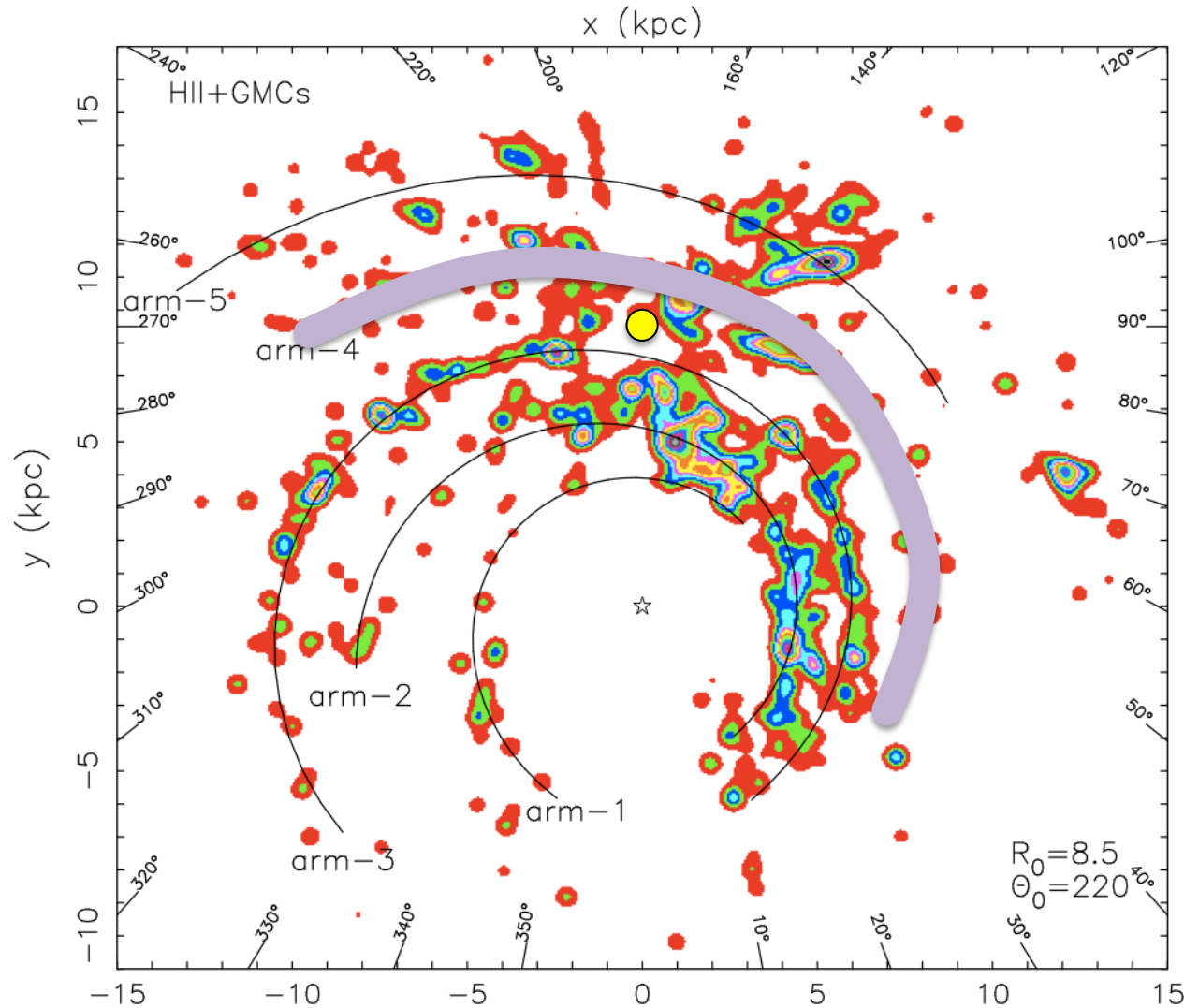
October 2009

Cloud Classification

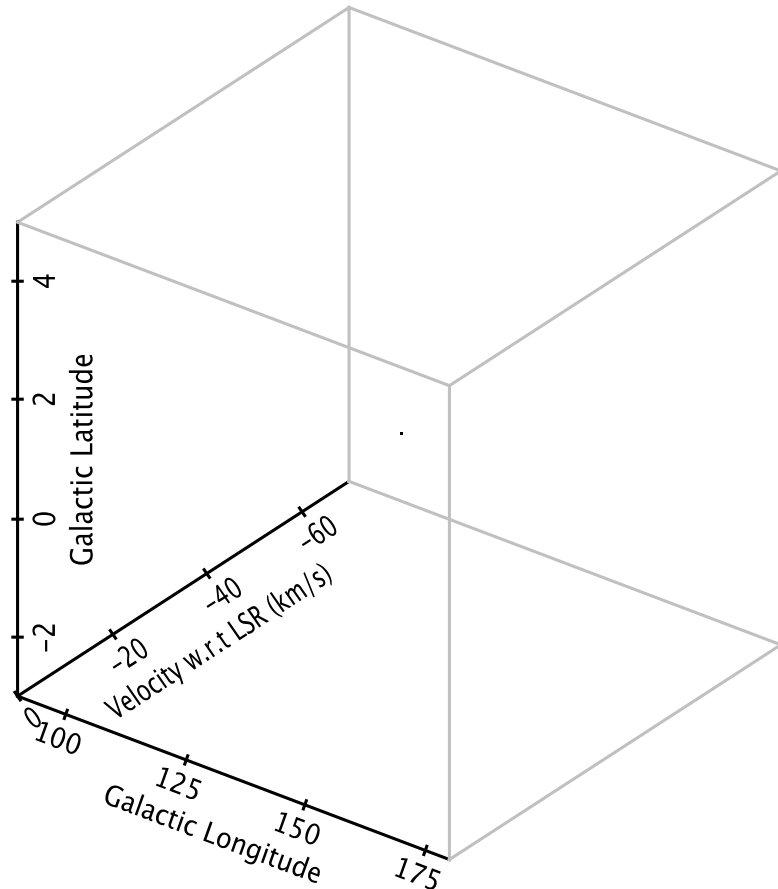


Credit – Leisawitz et al. 1989

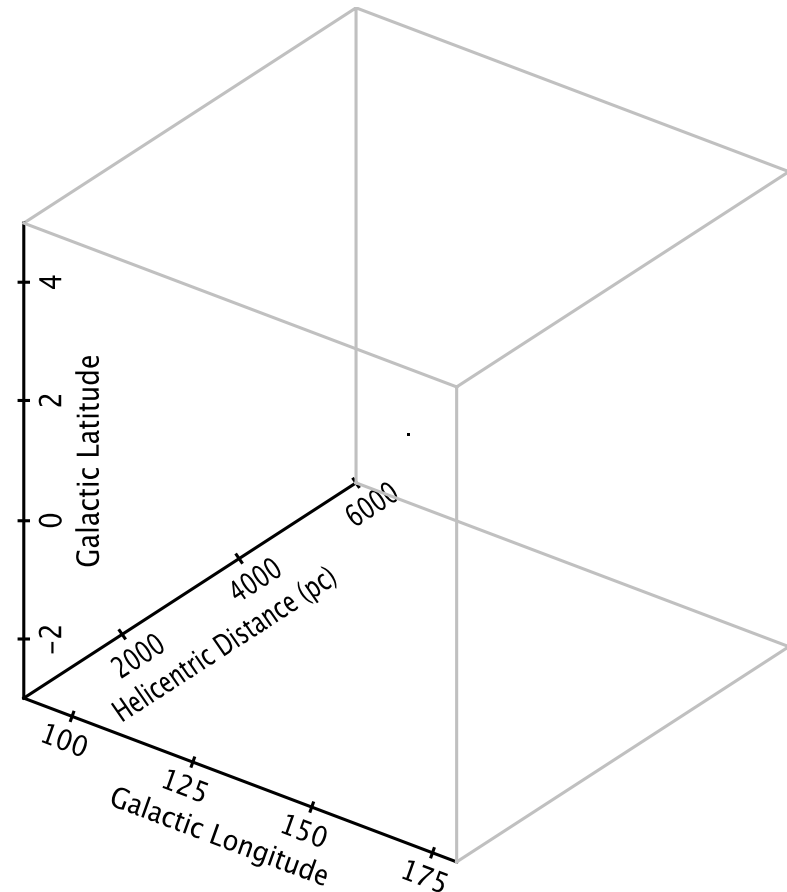
Distribution of Gases in the Milky Way



Parameter Space



$$(\theta_x, \theta_y, \frac{dz}{dt})$$

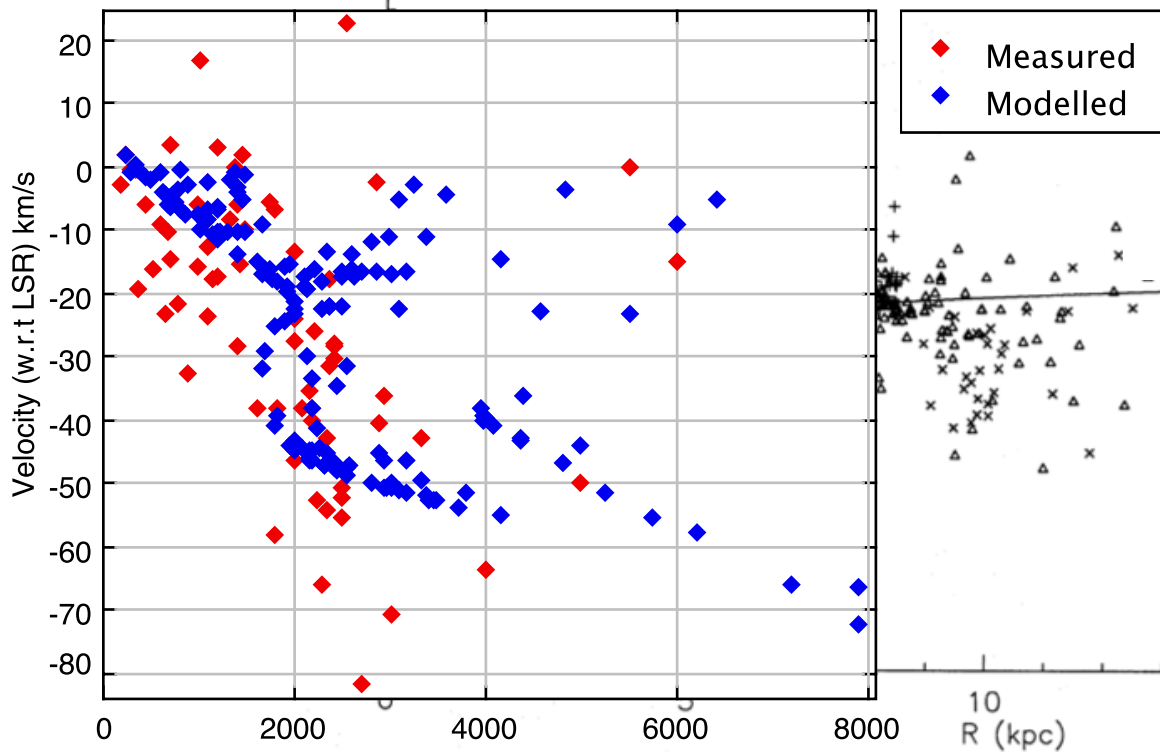


$$(\theta_x, \theta_y, z)$$

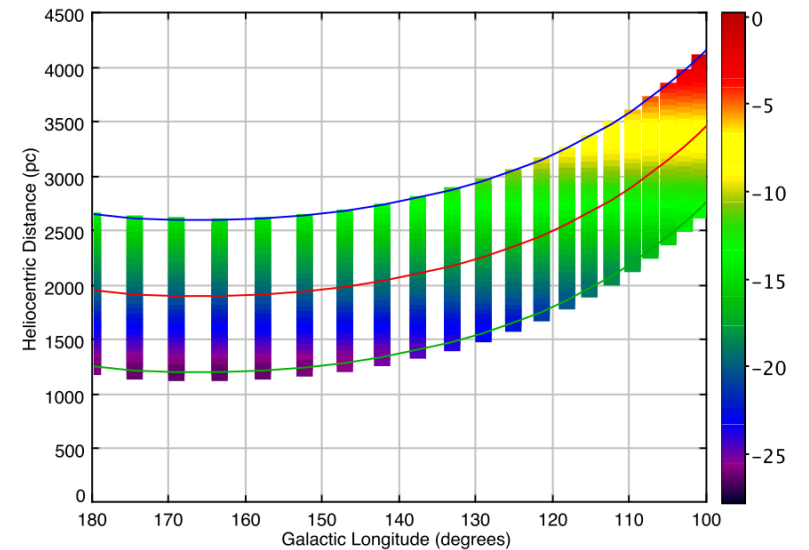
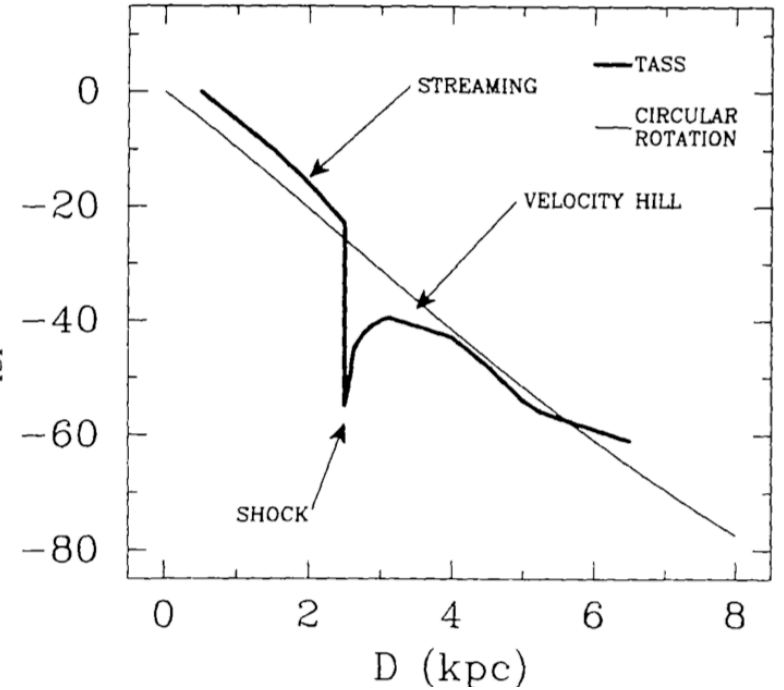
Velocity Field

Potential \rightarrow Velocity Shift

$$\Delta V_{LSR}^{shock} = \left(8.74 \pm 1.10 \cdot \frac{R}{[\text{kpc}]} \right) - (111.09 \pm 13.34)$$



Lee Summers Heliocentric Distance (pc)

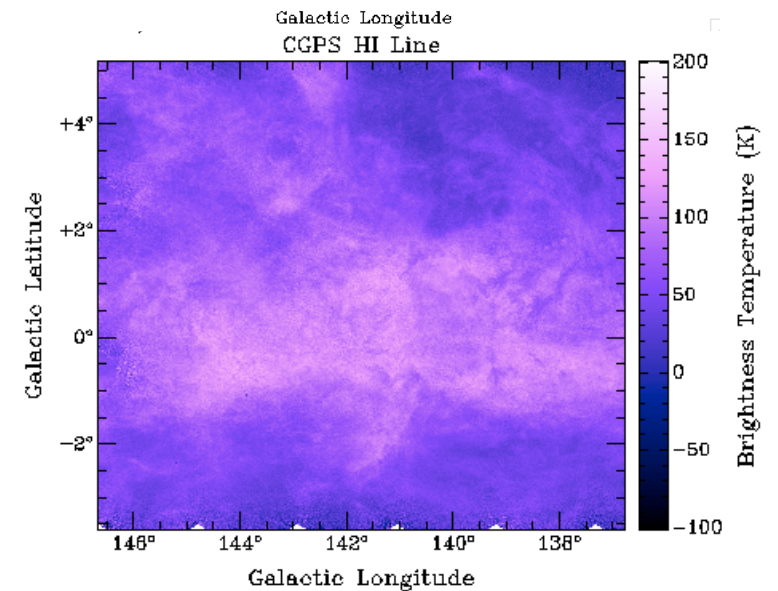
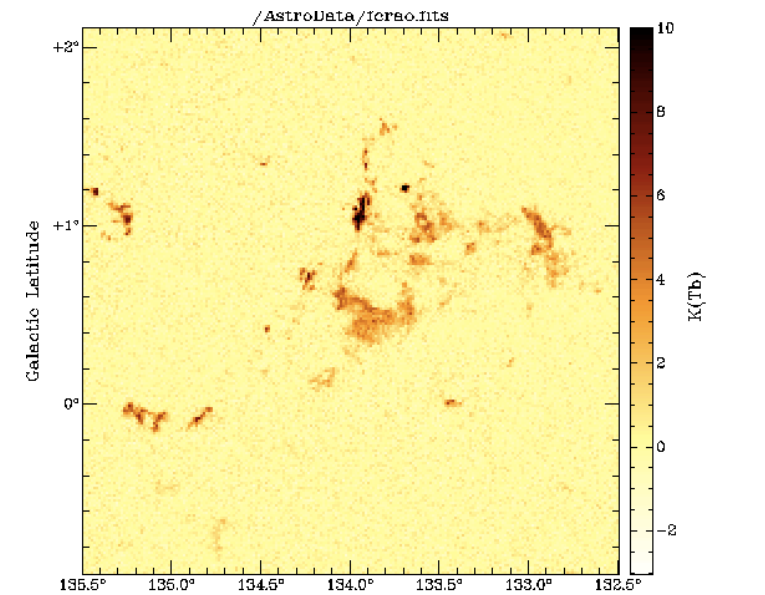


Mass Calculation

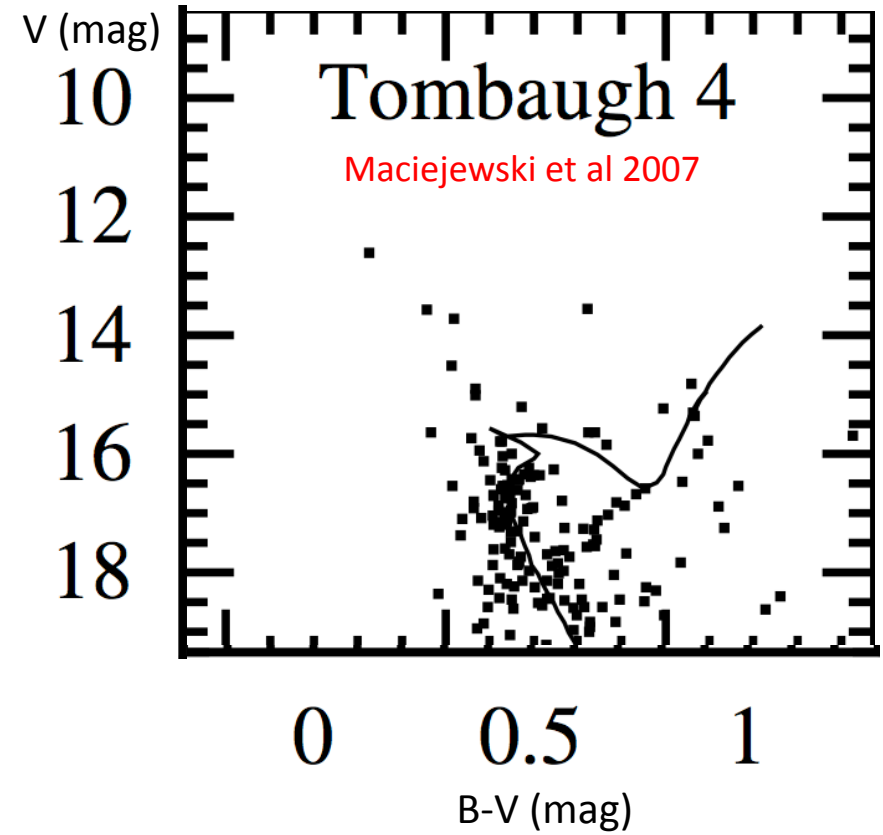
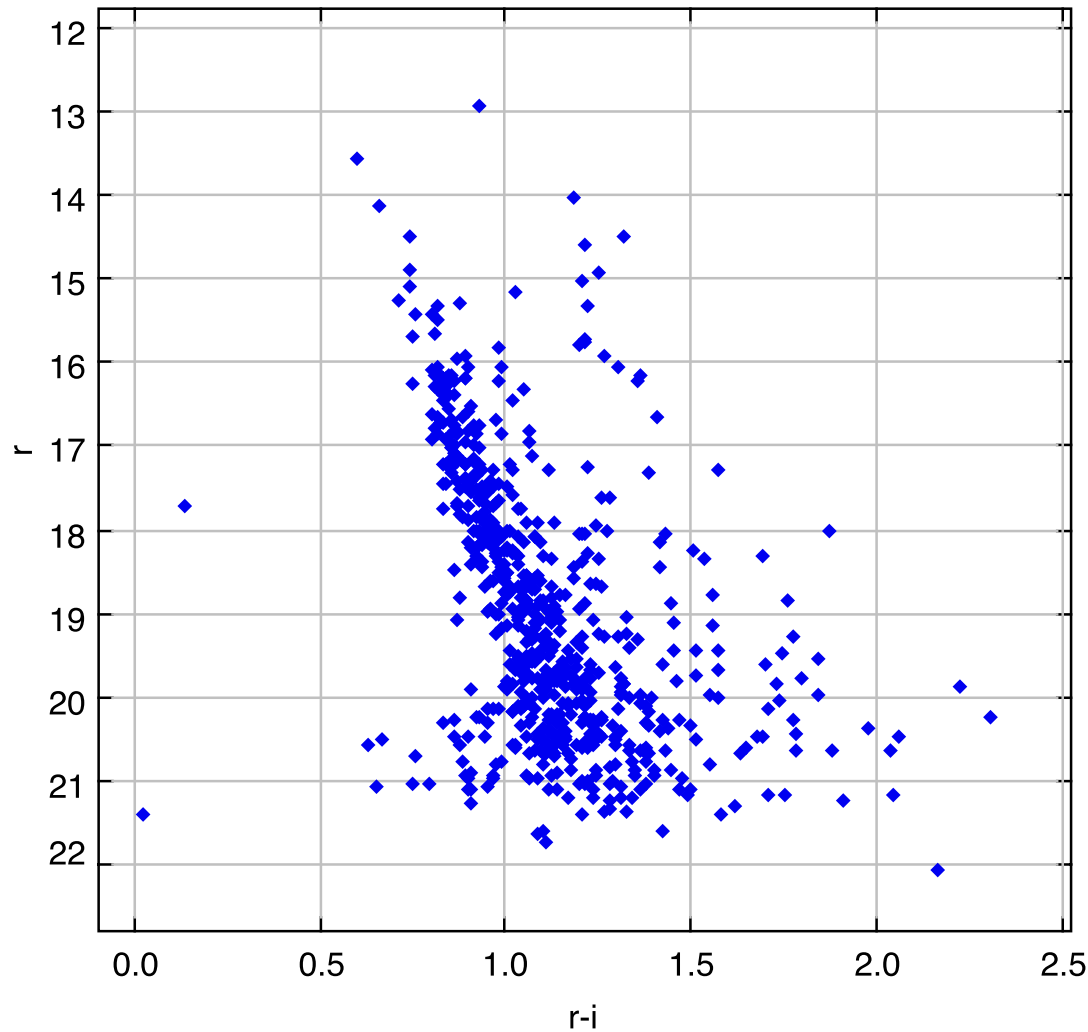
$$W = \int_{r=0}^r \int_{-\delta v}^{+\delta v} T \cdot dV \cdot 2\pi r dr$$

$$N_{\text{col}} = X \cdot W$$

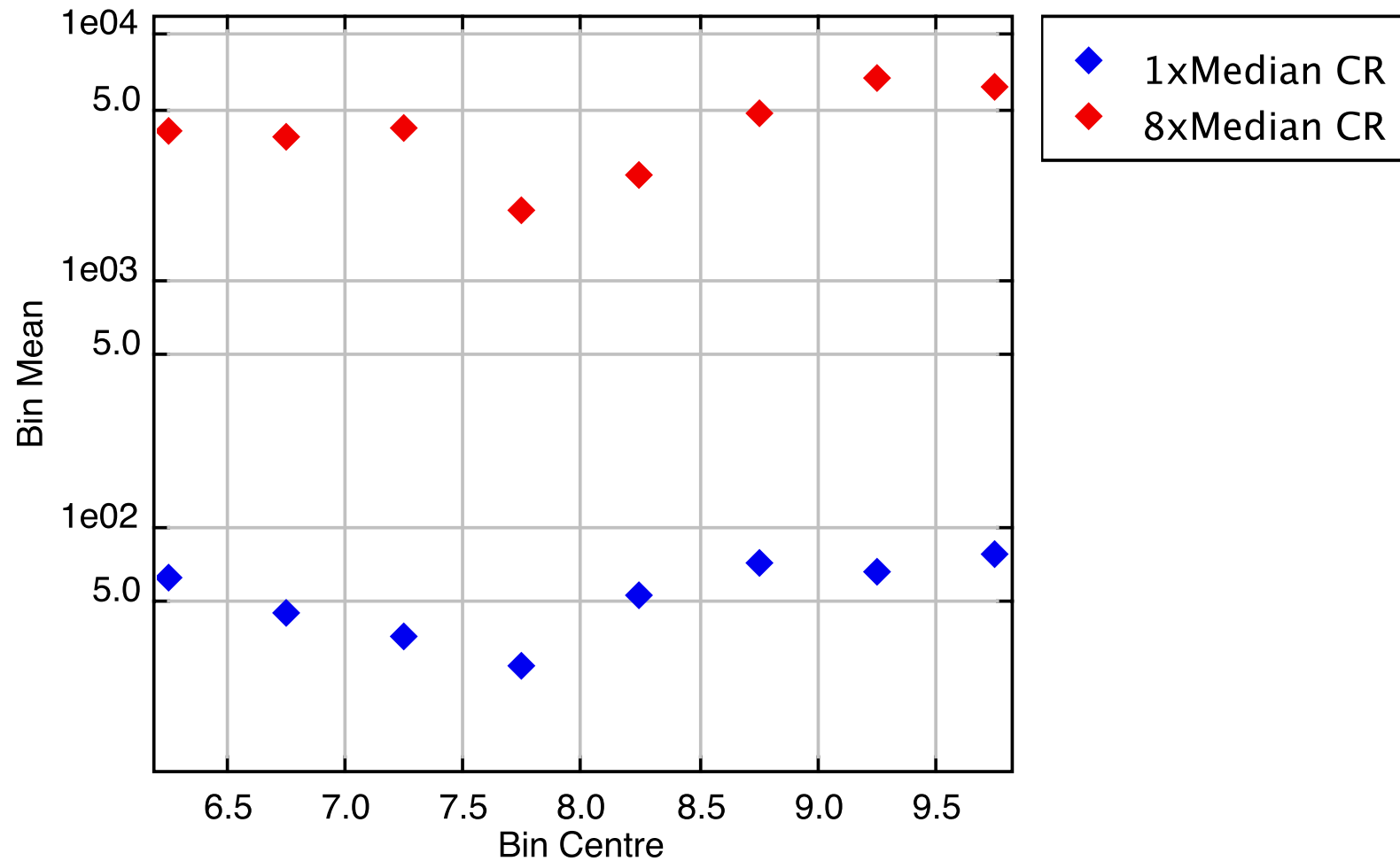
$$M_{\text{tot}} = N_{\text{col}} \cdot \text{Area} \cdot M_{\text{molecule}}$$



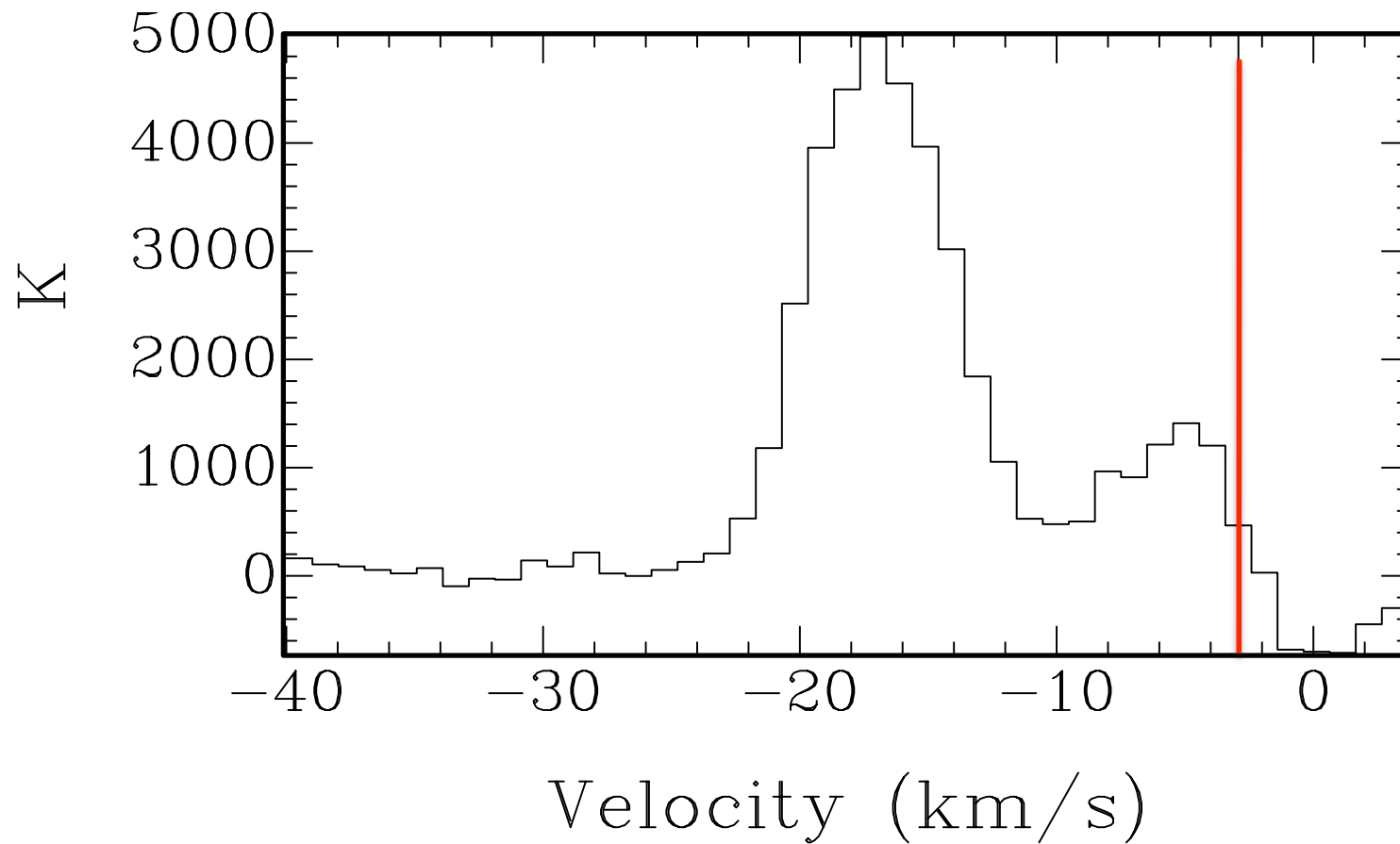
Catalogue Discrepancies



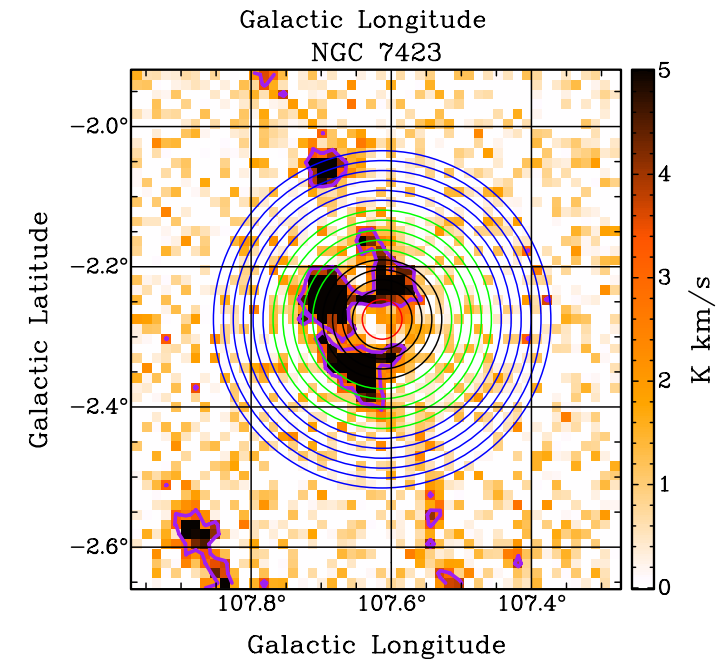
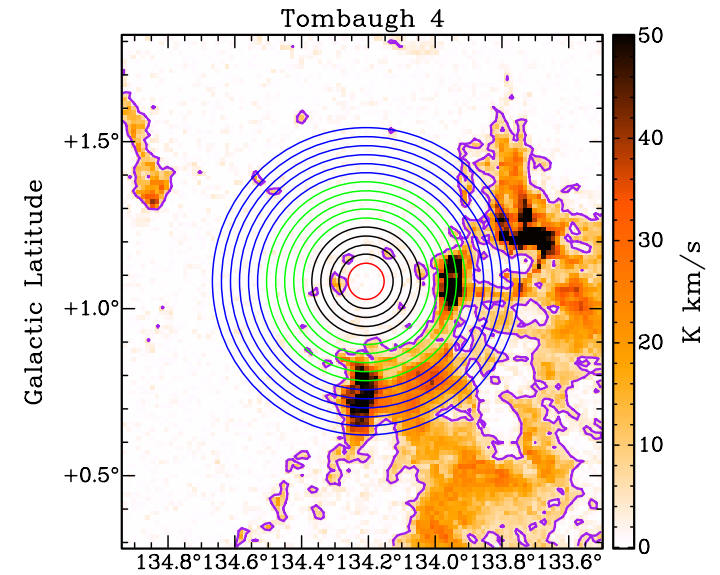
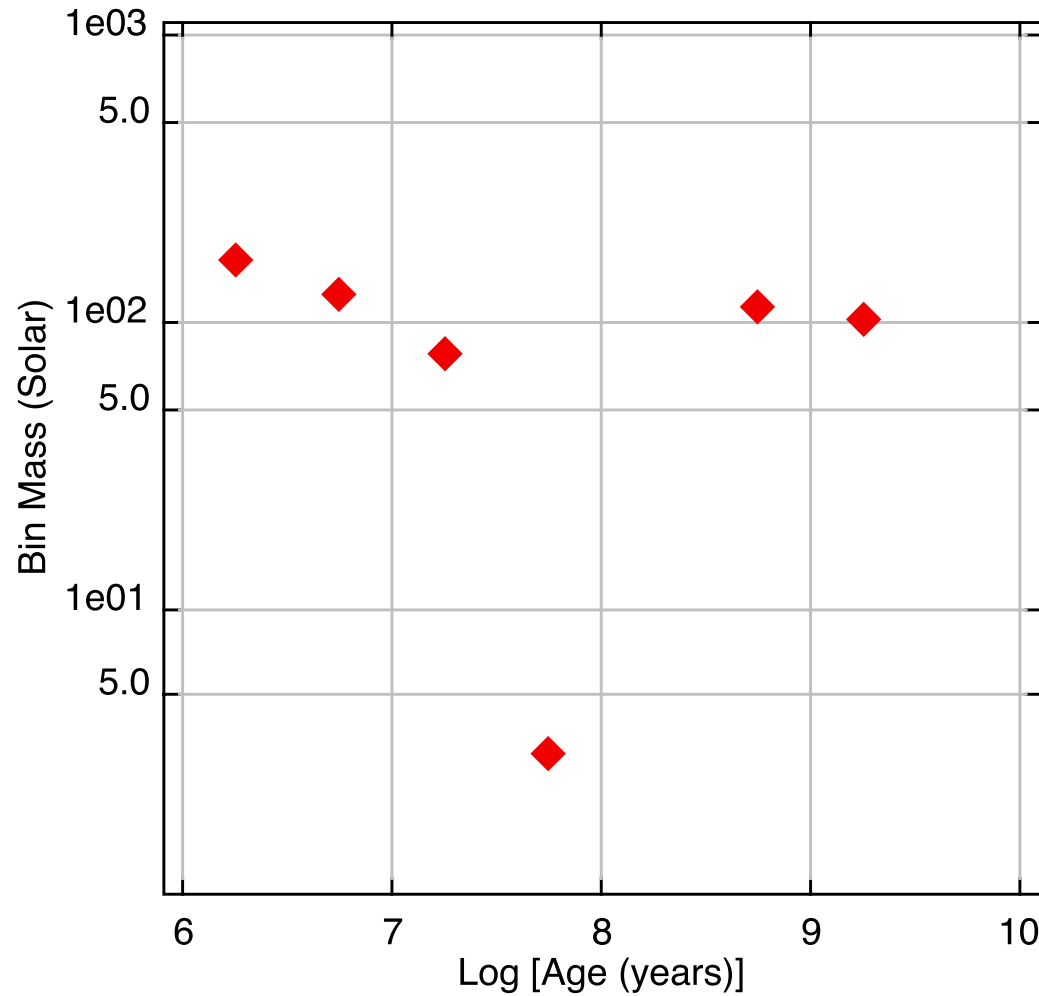
Preliminary Results



Modification of the Method



Associated Mass



Summary

- Physical association of Gas $< 10\text{pc}$
- Residual gas of formation $< 10^8\text{ yrs}$
- Objectively, material is invariant with age of cluster

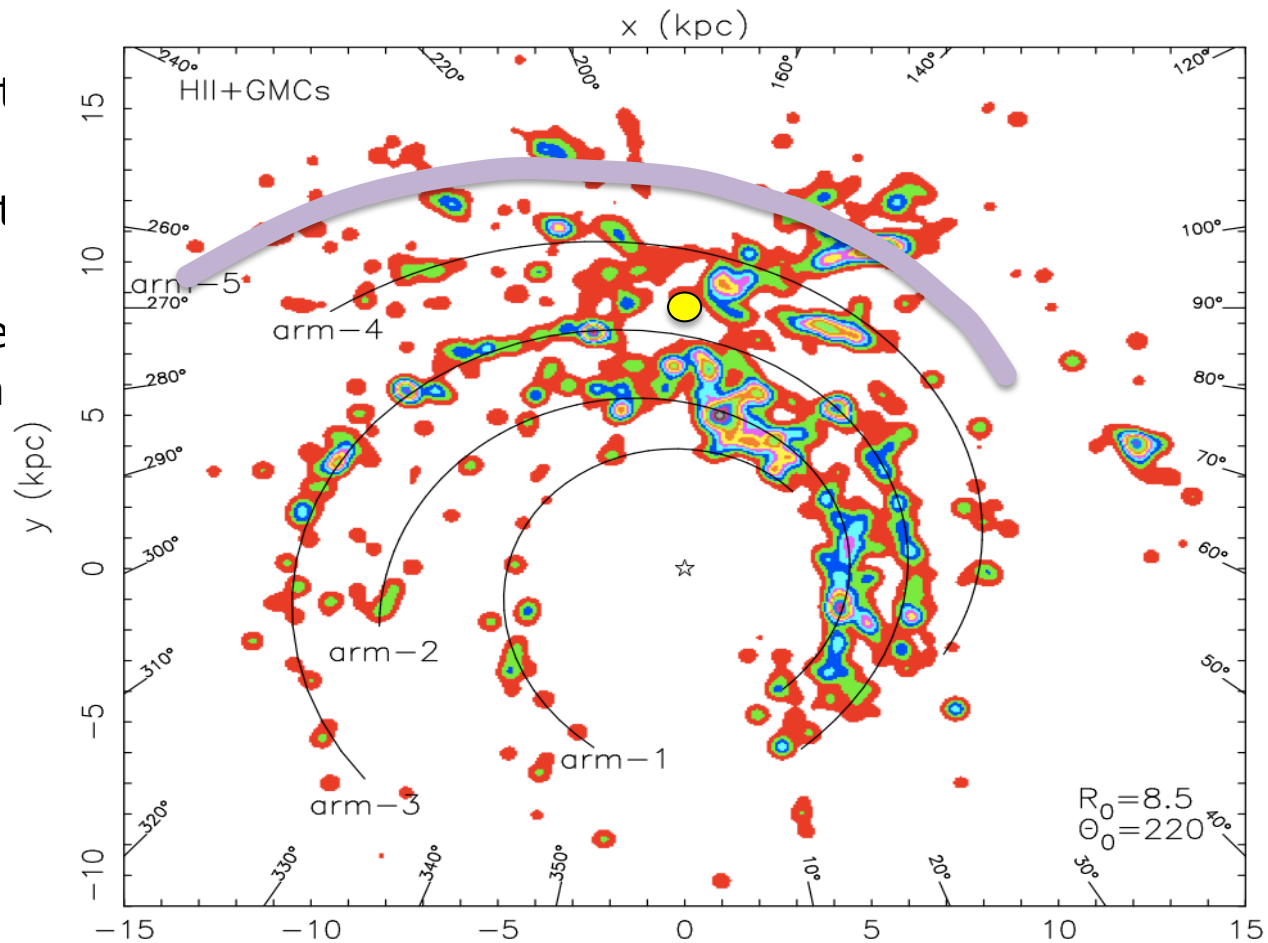
Further Work

- Mapping the cluster/gas associations of the outer spiral arm

- Refine t

- Charact

- High ve
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Thank You

Any Questions?